HIGH ROVIBRATIONAL STATES OF HC₃N: ANALYSIS OF THE RESONANCE NETWORK INVOLVING THE STRONG $2\nu_5$ OVERTONE BAND

L. BIZZOCCHI, M. MELOSSO, C. PUZZARINI, Dipartimento di Chimica "Giacomo Ciamician", via F. Selmi 2, 40126 Bologna (Italy); F. TAMASSIA, Dipartimento di Chimica Industriale "Toso Montanari", viale del Risorgimento 4, 40136 Bologna (Italy); O. PIRALI, M.-A. MARTIN DRUMEL, Université Paris-Saclay, CNRS, Institut des Sciences Moléculaires d'Orsay, 91405 Orsay (France)

Cyanoacetylene (HC₃N) is a major astrophysical tracer which manifests its presence in Space via strong rotational and infrared emissions. One of the most conspicuous features of its infrared (IR) spectrum is the $2\nu_5$ overtone bands located at \sim 1300 cm⁻¹ (7.7 μ m), and falling near the edge of mid infrared atmospheric window. Despite extensive past investigation of the HC_3N rovibrational spectrum [1, 2], this band still lacks an accurate analysis which takes into account the complex resonance network involving the $v_5 = 2$, bending state. Here, we report on a comprehensive study of the third polyad of interacting lelvels which includes the states: $(v_4 = 1, v_7 = 2), v_6 = v_7 = 2, (v_5 = 1, v_7 = 3), v_7 = 6, and v_5 = 2, located around$ 1300 cm^{-1} . The analysis is performed on an extensive dataset collected at SOLEIL (FAR/mid-IR) and at the University of Bologna (mid-IR and rotational data). Our objective is to retrieve a set of spectroscopic constants which have both clear physical meaning and excellent spectral predictive capability. To this aim, we adopt the ro-vibrational Hamiltonian already used for the analysis of the first and second resonance systems of HC_3N (e.g. [2]). It includes explicitly vibrational and ro-vibrational *l*-type resonance effects in bending overtones/combinations as well as the numerous anharmonic resonances occurring between accidentally nearly-degenerate states. The results will serve to improve the data already included in the 2020 release of the HI-TRAN database.

References

- [1] L. Bizzocchi, F. Tamassia, M. Melosso, et al., Astrophys. J. Suppl. S. 233, 11 (2017).
- [2] F. Tamassia, L. Bizzocchi M. Melosso, et al., J. Quant. Spectrosc. Rad. Trans. 279, 108044 (2021).