

## VUV-FT SPECTROSCOPY OF THE $a^3\Pi - X^1\Sigma^+$ CAMERON BAND SYSTEM IN THE CO MOLECULE

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The study of carbon monoxide is of considerable importance in modelling the evolution of nebulae, stars, planets and exoplanets, as well as entire galaxies, due to the second-largest abundance of this molecule in the universe after the molecular hydrogen.<sup>1</sup> Cameron system,  $a^3\Pi - X^1\Sigma^+$  is the subject of very intense study to understand a daylight of the Mars atmosphere.<sup>2,3,4</sup>

This work is aimed to obtain the highest energy, non-perturbed band among all the hot bands of the  $a^3\Pi - X^1\Sigma^+$  Cameron system in the CO molecule, namely the (3, 0), as well as to precisely analyze the  $a^3\Pi(v=3)$  vibrational level. For this purpose, the VUV-FT absorption technique was used by employing the wave-front-division spectrometer working as the end station on the DESIRS beamline (SOLEIL synchrotron).<sup>5,6</sup> The wavenumbers of the 9 rotational branches of the band under consideration were thus obtained. Based on the effective Hamiltonian and using the PGOPHER program<sup>7</sup>, the accurate rotational and  $\Lambda$ -doubling molecular constants as well as spin-orbit, spin-rotation, and off-diagonal spin-spin coupling constants of the  $a^3\Pi(v=3)$  level were obtained. The ro-vibronic terms of  $a^3\Pi(v=3)$ , based on the experimental transition frequencies, were calculated.

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<sup>1</sup>[doi:10.1146/annurev.astro.42.053102.134010](https://doi.org/10.1146/annurev.astro.42.053102.134010), E. F. Van Dishesock, *Annual Review of Astronomy and Astrophysics*, **42**, 119-167, (2004).

<sup>2</sup>[doi:10.1016/j.pss.2011.08.010](https://doi.org/10.1016/j.pss.2011.08.010), S.K. Jain *et al.*, *Planetary and Space Science*, **63-64**, 110-112 (2012).

<sup>3</sup>[doi:10.1016/j.icarus.2014.09.051](https://doi.org/10.1016/j.icarus.2014.09.051), A. Stiepen *et al.*, *Icarus*, **245**, 295-305, (2015).

<sup>4</sup>[doi:10.1016/j.icarus.2015.09.023](https://doi.org/10.1016/j.icarus.2015.09.023), L. Soret *et al.*, *Icarus*, **264**, 398-406, (2016).

<sup>5</sup>[doi:10.1103/PhysRevA.84.052509](https://doi.org/10.1103/PhysRevA.84.052509), A. J. de Nijs *et al.*, *Physical Review A*, **84**, 052509, (2011).

<sup>6</sup>[doi:10.1107/S1600577516006135](https://doi.org/10.1107/S1600577516006135), N. de Oliveira *et al.*, *Journal of Synchrotron Radiation*, **23**, 887-900, (2016).

<sup>7</sup>[doi:10.1016/j.jqsrt.2016.04.010](https://doi.org/10.1016/j.jqsrt.2016.04.010), C. Western, *Journal of Quantitative Spectroscopy and Radiative Transfer*, **186**, 221-242, (2017).